

# MOVING ENVELOPES OF STARS



V. V. Sobolev

Translated by Sergei Gaposchkin

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# Moving Envelopes Of Stars

**V. V. Sobolev**



## Moving Envelopes Of Stars:

Moving Envelopes of Stars Viktor Viktorovič Sobolev, 1955

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**The Astrophysics of Emission-Line Stars**

Tomokazu Kogure, Kam-Ching Leung, 2010-05-05 Many types of stars show conspicuous emission lines in their optical spectra. These stars are broadly referred to as emission line stars. Emission line stars are attractive to many people because of their spectacular phenomena and their amazing varieties and variability. This book offers general information on emission line stars starting from a brief introduction to stellar astrophysics and then moving to a broad overview of emission line stars including early and late type stars as well as pre main sequence stars. Detailed references and index are prepared for further reading. Tomokazu Kogure's email address: tkogure.pa2@so-net.ne.jp

**NASA Technical Translation**, 1969

*Luminous*

*Stars and Associations in Galaxies* International Astronomical Union. Symposium, 1986-06-30. Proceedings of the 116th

Symposium of the International Astronomical Union held at Porto Heli, Greece, May 26-31, 1985

*New Perspectives on*

*Stellar Pulsation and Pulsating Variable Stars* James M. Nemec, Jaymie M. Matthews, 1993-10-21

How can the interior of the Sun, white dwarfs and other stars be studied by stellar seismology? What can Doppler imaging tell us about high degree pulsations? What impact are CCD and infrared observations having on extending the Cepheid and RR Lyrae distance scale?

And how are other classes of pulsators providing independent checks of the distance scale? These and many other critical questions are answered in this timely review of the dramatic advances made in pulsating star research in the last decade. This survey collects together more than thirty comprehensive reviews and over one hundred summaries of research papers from the 139th IAU Colloquium held in Victoria, British Columbia. Together these cover all aspects of recent developments in the field of variable star research and preview some of the exciting advances anticipated for the next decade. This volume provides an essential review for graduate students and researchers.

Nuclear Science Abstracts, 1976-05

**Physics of**

**Thermal Gaseous Nebulae** L.H. Aller, 2012-12-06

Gaseous nebulae offer outstanding opportunities to atomic physicists, spectroscopists, plasma experts and to observers and theoreticians alike for the study of attenuated ionized gases. These

nebulae are often dusty, heated by radiation fields and by shocks. They are short-lived phenomena on the scale of a stellar lifetime, but their chemical compositions and internal kinematics may give important clues to advanced stages of stellar

evolution. The material herein presented is based on lectures given at the University of Michigan, University of Queensland,

University of California, Los Angeles and in more abbreviated form at the Raman Institute at the Scuola Internazionale di

Trieste and elsewhere. Much of it is derived originally from the series *Physical Processes in Gaseous Nebulae* initiated at the

Harvard College Observatory in the late 1930s. I have tried to emphasize the basic physics of the mechanisms involved and

mention some of the uncertainties that underlie calculations of many basic parameters. Emphasis is placed on ionized

plasmas with electron temperatures typically in the neighborhood of 10 000 K. Dust and other ingredients of the cold component of the interstellar medium are treated briefly from the point of view of their relation to hot plasmas of H II regions and planetary nebulae. Chemical composition determinations for nebulae are discussed in some detail while the last section deals with interpretations of elemental abundances in the framework of stellar evolution and nucleogenesis. Gaseous nebulae offer some particularly engaging opportunities for studies of stellar evolution. Physics of Stars and Stellar Systems Aleksandr Aleksandrovich Mikhailov, 1969 **Accretion Processes in Star Formation** Lee Hartmann, 2000-11-15 This first comprehensive account of the dynamical processes in the formation of stars and disks from which planets ultimately form

**Stellar Physics** Gennady S. Bisnovatyi-Kogan, 2011-01-04 Stellar Physics is an outstanding book in the growing body of literature on star formation and evolution. Not only does the author, a leading expert in the field, very thoroughly present the current state of knowledge on stellar physics but he handles with equal care the many problems that this field of research still faces. A bibliography with well over 1000 entries makes this book an unparalleled reference source. Stellar Evolution and Stability is the second of two volumes and can be read as can the first volume Fundamental Concepts and Stellar Equilibrium as a largely independent work. It traces in great detail the evolution of protostars towards the main sequence and beyond this to the last stage of stellar evolution with the corresponding vast range from white dwarfs to supernovae, explosions, gamma ray bursts and black hole formation. The book concludes with special chapters on the dynamical, thermal and pulsing stability of stars. This second edition is carefully updated in the areas of pre-supernova models, magnetorotational supernovae and the theory of accretion disks around black holes. Additional sections have been added on strange quark stars, jet formation and collimation, radiation driven winds in strong gravitational fields and gamma ray bursts. **Scientific American**, 1908

*The Brightest Stars* C. de Jager, 2012-12-06 No part of the Hertzsprung-Russell diagram shows a more pronounced diversity of stellar types than the upper part which contains the most luminous stars. Can one visualize a larger difference than between a luminous young and extremely hot O star and a cool evolved pulsating giant of the Mira type or an S type supergiant or again at the other side of the diagram the compact nucleus of a planetary nebula? But there is order and unity in this apparent disorder. Virtually all types of bright stars are evolutionally related in one way or the other. Evolution links bright stars. In many cases the evolution is speeded up by or at least intimately related to various signs of stellar instability. Bright stars lose mass either continuously or in dramatic sudden events; they vibrate or pulsate and with these tenuous gigantic objects this often happens in a most bizarre fashion. Sometimes the evolution goes so fast that fundamental changes are observable in the time span of a human's life; several of such cases have now been identified. **The Physics of Stars** Mathias Scholz, 2025-04-30 Interested students in the natural and engineering sciences as well as high school graduates, instructors, teachers and amateur astronomers will find a valuable overview of the physics of stars in this book. The only prerequisite is a basic mathematical and physical background which does not go beyond the knowledge of integral and

differential calculus In this regard this book aims to bridge the gap with the specialized literature available on the internet allowing readers to benefit from it The first part traces the historical development that led to a detailed understanding of the nature of stars and their life cycles The goal of the following chapters is to provide a pragmatic introduction to the physical processes that determine the structure and evolution of stars based on their fundamental parameters such as mass and chemical composition It will show what can be learned from the analysis of starlight about stellar atmospheres the fundamental role of the virial theorem in the lives of stars and the nuclear processes deep inside stars that provide the energy that makes them shine Finally there will be an in depth phenomenological look at the final stages of stellar evolution This section will discuss states of matter that are far from experimental realization but whose properties can be at least in principle inferred from the observation of concrete objects such as white dwarfs or neutron stars Exciting developments are still expected in this area in the future Mathias Scholz is hobby astronomer He studied physics at the University of Rostock from 1981 to 1986 Interested students in the natural and engineering sciences as well as high school graduates instructors teachers and amateur astronomers will find a valuable overview of the physics of stars in this book The only prerequisite is a basic mathematical and physical background which does not go beyond the knowledge of integral and differential calculus In this regard this book aims to bridge the gap with the specialized literature available on the internet allowing readers to benefit from it The first part traces the historical development that led to a detailed understanding of the nature of stars and their life cycles The goal of the following chapters is to provide a pragmatic introduction to the physical processes that determine the structure and evolution of stars based on their fundamental parameters such as mass and chemical composition It will show what can be learned from the analysis of starlight about stellar atmospheres the fundamental role of the virial theorem in the lives of stars and the nuclear processes deep inside stars that provide the energy that makes them shine Finally there will be an in depth phenomenological look at the final stages of stellar evolution This section will discuss states of matter that are far from experimental realization but whose properties can be at least in principle inferred from the observation of concrete objects such as white dwarfs or neutron stars Exciting developments are still expected in this area in the future

*Course in Theoretical Astrophysics* Viktor Viktorovich Sobolev, 1969      *The First Galaxies* Tommy Wiklund, Bahram Mobasher, Volker Bromm, 2012-12-15 New observations of the period between the cosmic recombination and the end of reionization are posing intriguing questions about where the first generations of stars were formed how the first galaxies were assembled whether these galaxies have low redshift counterparts and what role the early galaxies played in the reionization process Combining the new observational data with theoretical models can shed new light on open issues regarding the star formation process its role in the reionization of the Universe and the metal enrichment in galaxies at those early epochs This volume brings together leading experts in the field to discuss our current level of understanding and what may come in the near future as our observational as well as theoretical tools improve The book confronts the theory of how

the first stars black holes and galaxies formed with current and planned observations This synthesis is very timely just ahead of the establishment of major new facilities such as the James Webb Space Telescope JWST a next generation millimeter sub millimeter observatory in the Atacama desert ALMA and ground based Extremely Large Telescopes ELT Together they will revolutionize the study of the most distant objects in the Universe This volume is aimed at beginning graduate students but can also serve as a reference work for active researchers in the field Apart from presenting the fundamental concepts involved it also provides an introduction to the methods and techniques used The book will also be useful to anyone with an astrophysical background who needs an effective starting point for learning about the first stars and galaxies

*An Introduction to Astrophysical Hydrodynamics* Steven N. Shore, 2012-12-02 This book is an introduction to astrophysical hydrodynamics for both astronomy and physics students It provides a comprehensive and unified view of the general problems associated with fluids in a cosmic context with a discussion of fluid dynamics and plasma physics It is the only book on hydrodynamics that addresses the astrophysical context Researchers and students will find this work to be an exceptional reference Contents include chapters on irrotational and rotational flows turbulence magnetohydrodynamics and instabilities

**The Role of Dust in the Formation of Stars** Hans U. Käufl, Ralf Siebenmorgen, 2013-06-29 This ESO workshop which took place in September 1995 on a topic that at a first glance could be considered rather specialized attracted an unpredictably large number of scientists This certainly reflects the importance of this field which has lost its seemingly esoteric character in a wider astro physical context To give as much room as possible in these proceedings to the targeted talks no presentation of the Very Large Telescope Observatory has been included All readers missing such a presentation are reminded that up to date in depth information about the VLT status is available electronically 1 Papers were given concerning observations in the entire electromagnetic spectrum from x rays to mm waves i e exceeding 22 octaves in frequency The VLT as any ground based optical observatory can only address at best 7 octaves Nevertheless the VLT most likely the only ground based observatory specifically designed to access all these 7 octaves of the electromagnetic spectrum practically in parallel will undoubtedly be a tool of extreme value to this field

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