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Radiation Hydrodynamics in Stars and Compact Objects

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Radiation Hydrodynamics In Stars And Compact Objects

Dimitri Mihalas



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Radiation Hydrodynamics in Stars and Compact Objects Dimitri Mihalas, 1986 Radiation Hydrodynamics in Stars and Compact Objects International Astronomical Union. Colloquium, 1986 **Astronomy and Astrophysics Abstracts** S. Böhme, U. Esser, H. Hefele, I. Heinrich, W. Hofmann, D. Krahn, V. R. Matas, L. D. Schmadel, G. Zech, 2013-12-14 From the reviews Astronomy and Astrophysics Abstracts has appeared in semi annual volumes since 1969 and it has already become one of the fundamental publications in the fields of astronomy astrophysics and neighbouring sciences It is the most important English language abstracting journal in the mentioned branches The abstracts are classified under more than hundred subject categories thus permitting a quick survey of the whole extended material The AAA is a valuable and important publication for all students and scientists working in the fields of astronomy and related sciences As such it represents a necessary ingredient of any astronomical library all over the world Space Science Reviews 1 Dividing the whole field plus related subjects into 108 categories each work is numbered and most are accompanied by brief abstracts Fairly comprehensive cross referencing links relevant papers to more than one category and exhaustive author and subject indices are to be found at the back making the catalogues easy to use The series appears to be so complete in its coverage and always less than a year out of date that I shall certainly have to make a little more space on those shelves for future volumes The Observatory Magazine 1 **Physical Processes in Comets, Stars and Active Galaxies** Wolfgang Hillebrandt, Emmi Meyer-Hofmeister, Hans-Christoph Thomas, 2012-12-06 In May 1986 a two day workshop on Physical Processes in Comets Stars and Active Galaxies was held at the Ringberg Castle near Lake Tegernsee and this rather unusual collection of topics needs a few words of explanation When we first thought of organizing a workshop on such a large variety of astrophysical objects our main motivation was to honor Rudolf Kippenhahn and Hermann Ulrich Schmidt on the occasion of their 60th birthdays and we planned to cover at least a fraction of their fields of active research We then realized immediately that despite the fact that the objects are so different the physical processes involved are very much the same and that it is this aspect of astrophysics which governed the scientific lives of both of our distinguished colleagues and friends and allowed them to make major contributions to all those fields Apparently this viewpoint was shared by many colleagues and it was therefore not surprising that in response to our invitation everybody who had been invited agreed to come and to present a talk The workshop then turned out to be a real success In contrast to highly specialized conferences fundamental problems as well as very recent developments were discussed and the participants appreciated the opportunity to exchange ideas Supernovae - Proceedings Of The 6th Jerusalem Winter School For Theoretical Physics Tsvi Piran, Steven Weinberg, J Craig Wheeler, 1990-07-23 Supernovae are highly energetic phenomena for which it is necessary to use simultaneously particle physics nuclear physics and hydrodynamics to study the creation of the strong explosions involved Supernovae synthesize heavy elements and in some cases lead to the formation of neutron stars or black holes Recent progress has

distance of only 800 pc Chevalier and Seward 1988 must have awed our ancestors Tycho's supernova of 1572 at a distance of 2500 pc had a magnitude of 4.0 comparable to Venus at its brightest and Kepler's supernova of 1604 had a magnitude of 3 or so Thus the Cygnus Loop supernova might have had a magnitude of 6 or so and should have been readily visible in daytime A supernova in Vela about 8000 B.C. was comparably close as was SN 1006 whose magnitude may have been 9 While most of the supernova records come from the Old World the supernova of 1054 is recorded in at least one petroglyph in the American West

Dynamical Spacetimes and Numerical Relativity Joan M. Centrella, 1986-09-18 *High-Energy Astrophysics Research Enabled By The Probe-Class Mission Concept HEX-P* Gabriele Ponti, Murray Brightman, Esra Bulbul, Renee Ludlam, Daniel Stern, Javier A. García, 2025-06-02 The Astro 2020 Decadal report Pathways to Discovery in Astronomy and Astrophysics for the 2020s identified the need for a new class of mission for the next decade that fills the gap between a Mid-sized Explorer MIDEX and Flagship Large mission The report specifically recommends that a cadence of one probe mission per decade with a cost cap of 1.5 billion balances scientific scope with timeliness NASA has followed this recommendation and opened a call for Astrophysics Probe Explorer APEX proposals restricted to either a far infrared or an X-ray mission due in Fall 2023 with a planned launch date in no later than 2032 In response to this call the High Energy X-ray Probe HEX-P concept has been developed as a partnership between NASA's Jet Propulsion Laboratory JPL the California Institute of Technology Caltech and the NASA Goddard Space Flight Center GSFC with international partnership from the Italian Space Agency ASI Italy the German Space Agency DLR and the Max Planck Institute for Extraterrestrial Physics MPE Germany HEX-P is a mission concept that offers sensitive broad band coverage from soft to hard X-rays 0.2-80 keV with an exceptional combination of spectral timing and angular capabilities It features two high energy telescopes HETs that focus hard X-rays and one low energy telescope LET that focuses lower energy X-rays building upon the heritage of previous successful missions such as XMM-Newton and NuSTAR With this leap in observational capability HEX-P will be capable of addressing fundamental questions about the extreme environments around black holes and neutron stars map the growth of supermassive black holes and quantify the effect they have on their environments HEX-P will resolve the hard X-ray emission from dense regions of our Galaxy to understand the high energy source populations and investigate dark matter candidate particles through their decay channel signatures In this Frontiers Research Topic we present a collection of papers which delve into some of the most exciting scientific questions that can be addressed with a mission like HEX-P The papers have been prepared by members of the HEX-P collaboration which is organized into four main scientific pillars Black Hole Growth Accretion Power Stellar Evolution Time Domain and Multi-Messenger Astronomy The individual papers cover a range of topics including The cosmic X-ray background Seyfert galaxies Compton-thick active galaxies and blazars Black hole binaries neutron stars magnetars ultraluminous X-ray sources and tidal disruption events The physics of the X-ray corona X-ray populations in nearby galaxies Supernova remnants pulsar wind nebulae and nuclear astrophysics Galactic PeVatrons star

clusters superbubbles microquasar jets and gamma ray binaries The Galactic Center Supermassive black spin measurements and dual active galaxies Pulsar Timing electromagnetic counterparts of gravitational wave sources and transient phenomena The main goal of this Research Topic is to present detailed modeling and simulations for a range of HEX P science cases in order to demonstrate the capabilities of HEX P in serving the astrophysics community in the next decade The members of the HEX P mission are welcome to present Original Research papers as well as any other kind of article types that fit their manuscript Please have a look at [here](#) to choose the appropriate article type to submit to this collection This Research Topic is organized by the researchers of the HEX P mission Authors and Guest Editors belong to the same consortium Furthermore Dr Daniel Stern and Dr Javier Garcia are the Principal Investigator and the Project Scientist of the mission respectively All manuscripts will be peer reviewed by researchers external to the collaboration

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