

because the data are so urgently needed for abundance determinations and for comparisons with other stellar spectra.

The range of the present table has been limited to that long span of the solar spectrum that overlaps the more recent and current observations from rocket and orbiting-solar-observatory spectra in the ultraviolet (short of 3000 Å), and from the Fraunhofer spectra in the infrared (long of 7500 Å). It is essentially a second revision of Rowland's Table corrected and supplemented by the Utrecht Atlas material. Even in this range the work is far from completed. Many more solar lines

will doubtless be added when the spectrum has been completely reobserved under the best observing conditions and with the best available modern instruments. The task of remeasuring the complex spectrum and of extending identifications and intensity measurements is a heavy one in itself, and one that requires, also, much additional work on the analysis of individual laboratory spectra. The present contribution is intended as a convenient handbook for the astrophysicist until the ideal self-contained compendium of the solar spectrum from the x-ray to the microwave region becomes a reality.

## 2. Description of the Solar Spectrum Compendium

The successive columns of the solar ledger are described. The general procedure by which the data on measured intensities have been obtained is fully explained. The many users require, also, more detailed information on the analyses of individual spectra than can be recorded here, for the many lines classified from laboratory investigations. Sufficient information is given in the ledger to provide convenient cross reference to the separate papers on spectrum analysis. For special details concerning individual lines the column headed "Notes" should be consulted.

### 2.1 The Wavelengths: Column One

A comprehensive revision of Rowland's *Preliminary Table of Solar Spectrum Wavelengths* entails a thorough study of the wavelengths of the lines reported by Rowland. These remarkably accurate measurements by Jewell, that date back to 1895-97 [2], were reduced to the International Scale of Wavelengths in the first revision of Rowland's Table in 1928 [3].

The tremendous amount of work that has gone into interferometric observations for precise wavelength standards need not be discussed here. Details may be found in the early Transactions of the International Astronomical Union [7], in the Allegheny Publications by Burns, Meggers, and Kloss [8], and in the Mount Wilson Contributions by St. John, H. D. Babcock and others. St. John discussed the method used for the reduction to the international system, and has published the correction curve based on the International Standard Wavelength Scale of 1922 [3]. The wavelengths listed in the 1928 edition are based on this curve. They are in three categories: (1) Lines marked "m". These lines were observed by Rowland, but remeasured for use in correcting Rowland's wavelength scale. (2) Lines marked "w", denoting weighted means of remeasured lines and Rowland's corrected values. (3) Undesignated wavelengths; these formed the great majority. They were lines listed by Rowland but corrected to the international scale of 1922 by means of the correction curve adopted in 1928.

Throughout the years in which the 1928 edition has been actively used, several problems regarding wavelengths have arisen. Further work on the International Standards has indicated that a small running correction is required to reduce the wavelengths based on the 1922 scale to the scale adopted in 1928 by the International Astronomical Union. These corrections, reported by H. D. Babcock [9] are as follows:

Table 1. Wavelength Corrections for Reduction to the 1928 Scale of Standards

Region (Å)	Correction to 1922 Scale (Å)	Region (Å)	Correction to 1922 Scale (Å)
2995-3133	-0.0012	6125-6200	-0.006
3133-3370	*(-0.0010)	6200-6455	-0.007
3370-3705	-0.0006	6455-6630	-0.008
3719-3849	-0.0015	6630	-0.009
3850-3969	-0.0015	6800	-0.010
4000-5000	-0.002	6900	-0.012
5000-5780	*(-0.003)	7000	-0.014
5780-5960	*(-0.004)	7100	-0.016
5960-6125	-0.005		

\* Values in parentheses are interpolated.

They must be applied to all wavelengths published in the 1928 edition, in order to comply with the standards recommended by Commission 14 of the Union.

Although Rowland's observations extend from 2975 Å to 7330 Å, his spectrograms did not record the solar spectrum completely in the ranges short of 3000 Å and long of 6600 Å. With the development of new photographic emulsions, Rowland's work has been superseded in the short- and long-wave regions. Meggers made a thorough study of the spectrum from 6500 Å to 9000 Å in 1919 [10] and by means of Doppler displacements, carefully separated the solar lines from the telluric lines. This work was subsequently extended at Mount Wilson by H. D. Babcock and others, by comparing line intensities at high and low sun. In the present table the

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**Scientific and Technical Aerospace Reports**, 1983      **The Solar Spectrum 3069Å-2095Å** Charlotte Emma Moore, 1982 This report presents the final listing of solar lines recorded in the NRL echelle spectra photographed at high resolution from Aerobee rockets flown in 1961 and 1964 The wavelength range covered is 3069 to 2095 It is intended to accompany NRL report numbers 7788 An Atlas of the Solar Spectrum Between 2226 and 2992 Angstroms which presents the solar irradiance at 0.03 resolution as derived from the echelle spectra Solar wavelengths are given to 0.01 and estimated solar intensities are listed on a visual scale of 1 to 9 Identifications have been made from a detailed study of the multiplets in individual spectra and are based on a search of the spectroscopic literature Laboratory wavelengths multiplet numbers and references are cited for each line Approximately 6150 lines are reported of which 80% are identified      **International Index to Periodicals**, 1929 An author and subject index to publications in fields of anthropology archaeology and classical studies economics folklore geography history language and literature music philosophy political science religion and theology sociology and theatre arts      *Astronomical Spectrographs and Their History* John Hearnshaw, 2009-02-12 Astronomical spectrographs analyse light emitted by the Sun stars galaxies and other objects in the Universe and have been used in astronomy since the early nineteenth century This book provides a comprehensive account of spectrographs from an historical perspective from their theory and development over the last two hundred years to the recent advances of the early twenty first century The author combines the theoretical principles behind astronomical spectrograph design with their historical development Spectrographs of all types are considered with prism grating or grism dispersing elements Included are Cassegrain coud prime focus chelle fibre fed ultraviolet nebular objective prism multi object instruments and those which are ground based on rockets and balloons or in space The book contains several tables listing the most significant instruments around 900 references and over 150 images making it an indispensable reference for professional astronomers graduate students advanced amateur astronomers and historians of science      Monthly Catalog of United States Government Publications, 1984      **Communications - Observatoire Royal de Belgique. Serie A.** Observatoire royal de Belgique, 1972      *Journal of Research of the National Bureau of Standards* United States. National Bureau of Standards, 1939      *The Solar Spectrum 2935 Å to 8770 Å* Charlotte Emma Moore, Marcel Gilles Jozef Minnaert, Jakob Houtgast, 1966      **The Analysis of Starlight** J. B. Hearnshaw, 1990-04-19 This book presents a detailed pedagogical account of the equation of state and its applications in several important and fast growing topics in theoretical physics chemistry and engineering This book is the story of the analysis of starlight by astronomical spectroscopy It describes the development of the subject from the time of Joseph Fraunhofer who in 1814 used a telescope mounted prism to observe the spectral light emitted from several bright stars He discovered that light was missing at certain colours wavelengths in the starlight and these so called spectral lines were subsequently shown to hold clues to the nature of the stars themselves The

book explains how the classification of stars using their line spectra developed into a major branch of astronomy whilst new methods in astrophysics made possible the approximate quantitative analysis of spectral lines in the 1920s and 1930s After the Second World War these techniques were considerably improved when computers were programmed to model the structure of the outer layers of stars Basic concepts in spectroscopy and spectral analysis are also covered and finally Dr Hearnshaw comments on the stellar spectroscopy of some individual star

**Bureau of Standards Journal of Research** United States. Bureau of Standards,1928 **Bureau of Standards Journal of Research** ,1931 *Communications to the National Academy of Sciences* Mount Wilson Observatory,1915\* **Bureau of Standards Journal of Research** United States. National Bureau of Standards,1933 **Publications of the Astronomical Society of the Pacific** Astronomical Society of the Pacific,1927 *Annals of the Cape Observatory* ,1911 *A STUDY OF THE SPECTRUM AND ORBIT OF UPSILON SAGITTARII*. FRANCES LOUISE SEYDEL,1929 **Einstein in Context** Mara Beller,Robert S. Cohen,Jürgen Renn,1993-10-29 This special issue of Science and Context examines the entire area of scientific inquiry surrounding Einstein presenting controversies and debates within their contexts *Publications of the Dominion Astrophysical Observatory* ,1971

Monthly Notices of the Royal Astronomical Society Royal Astronomical Society,1905 Portfolio of 8 charts accompanies v 83

**The Analysis of Starlight** John B. Hearnshaw,2014-03-17 First published in 1986 this is the story of the analysis of starlight by astronomical spectroscopy Beginning with Joseph Fraunhofer s discovery of spectral lines in the early nineteenth century this new edition continues the story through to the year 2000 In addition to the key discoveries it presents the cultural and social history of stellar astrophysics by introducing the leading astronomers and their struggles triumphs and disagreements Basic concepts in spectroscopy and spectral analysis are included so both observational and theoretical aspects are described in a non mathematical framework This new edition covers the final decades of the twentieth century with its major advances in stellar astrophysics the discovery of extrasolar planets new classes of stars and the observation of the ultraviolet spectra of stars from satellites The in depth coverage makes it essential reading for graduate students working in stellar spectroscopy professional and amateur astronomers and historians of science

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