

Numerical radiative transfer

- Real atoms have hundreds of different energy levels
 - Very computationally intensive
 - Many levels have little effect on the studied line
 - ⇒ Make simplified, smaller atomic models
- Compute one element at a time

Numerical Radiative Transfer

**Adrian Doicu, Thomas
Trautmann, Franz Schreier**



Numerical Radiative Transfer:

Numerical Methods in Multidimensional Radiative Transfer Guido Kanschat, Erik Meinköhn, Rolf Rannacher, Rainer Wehrse, 2008-12-24 Traditionally radiative transfer has been the domain of astrophysicists and climatologists In nuclear technology one has been dealing with the analogous equations of neutron transport In recent years applications of radiative transfer in combustion machine design and in medicine became more and more important In all these disciplines one uses the radiative transfer equation to model the formation of the radiation field and its propagation For slabs and spheres effective algorithms for the solution of the transfer equation have been available for quite some time In addition the analysis of the equation is quite well developed Unfortunately in many modern applications the approximation of a 1D geometry is no longer adequate and one has to consider the full 3D dependencies This makes the modeling immensely more intricate The main reasons for the difficulties result from the fact that not only the dimension of the geometric space has to be increased but one also has to employ two angle variables instead of one and very often one has to consider frequency coupling due to motion or redistribution in spectral lines In actual calculations this leads to extremely large matrices which in addition are usually badly conditioned and therefore require special care Analytical solutions are not available except for very special cases Although radiative transfer problems are interesting also from a mathematical point of view mathematicians have largely neglected the transfer equation for a long time

Numerical Radiative Transfer Wolfgang Kalkofen, 1988-01-07 First published in 1988 this book is a manual of methods for solving problems in radiative transfer Several of the methods on operator perturbation as well as on polarised radiative transfer appeared for the first time in this volume and the sections dealing with these topics each include introductory articles The methods described are principally directed at astrophysical plasmas but the methods are easily adaptable to applications involving other media where self absorption of the radiation is important The book is primarily intended for graduate students and research workers in the field of radiative transfer

Numerical Radiative Transfer with State-of-the-art Iterative Methods Made Easy, 2015

Radiative Heat Transfer Michael F. Modest, 2003-03-07 The basic physics of radiative heat how surfaces emit reflect and absorb waves and how that heat is distributed

Radiative Transfer Hélène Frisch, 2022-05-26 This book discusses analytic and asymptotic methods relevant to radiative transfer in dilute media such as stellar and planetary atmospheres Several methods providing exact expressions for the radiation field in a semi infinite atmosphere are described in detail and applied to unpolarized and polarized continuous spectra and spectral lines Among these methods the Wiener Hopf method introduced in 1931 for a stellar atmospheric problem is used today in fields such as solid mechanics diffraction theory or mathematical finance Asymptotic analyses are carried out on unpolarized and polarized radiative transfer equations and on a discrete time random walk Applicable when photons undergo a large number of scatterings they provide criteria to distinguish between large scale diffusive and non diffusive behaviors typical scales of variation of the radiation field such as the thermalization length and specific descriptions

for regions close and far from boundaries Its well organized synthetic view of exact and asymptotic methods of radiative transfer makes this book a valuable resource for both graduate students and professional scientists in astrophysics and beyond

Radiation Transfer Apresyan,2019-07-16 The authors expound on non traditional phenomena for transfer theory which are nevertheless of considerable interest in wave measurements and bring the advances of transfer theory as close as possible to the practical needs of those working in all areas of wave physics The book opens with a historical overview of the topic then moves on to examine the phenomenological theory of radiative transport blending traditional theory with original ideas The transport equation is derived from first principles and the ensuing discussion of the diffraction content of the transport equation and non classical radiometry is illustrated by practical examples from various fields of physics Popular techniques of solving the transport equation are discussed paying particular attention to wave physics and computing the coherence function The book also examines various problems which are no longer covered by the traditional radiative transfer theory such as enhanced backscattering and weak localization phenomena nonlinear transport problems and kinetic equations for waves This monograph bridges the gap between the simple power balance description in radiative transfer theory and modern coherence theory It will be of interest to researchers and professionals working across a wide range of fields from optics acoustics and radar theory to astrophysics radioastronomy and remote sensing as well as to students in these areas

Terrestrial Radiative Transfer Harriet H. Natsuyama,Sueo Ueno,Alan P. Wang,2012-12-06 In this book we share our work with those who are faced with the challenging problem of studying the earth s atmosphere and the interactions between the atmosphere and the earth s surface While there are some excellent books on this topic written from the physical point of view those discussing the modeling and computational aspects are few and far between Our book is intended to bridge this gap so that students as well as investigators will be able to understand and apply practical ways of determining solutions Radiative transfer theory on which this book is based is elegant and great minds have contributed to its richness Instead of duplicating the clas sical references we have taken a different approach We have developed the invariant imbedding approach both analytically and computationally because of its attractiveness for producing numerical solutions Having witnessed the transition to the computer age we know that a new attitude to mathematical formulation is required The one that we endorse is a model stated in the form of a Cauchy problem a system of ordinary differential equations with a complete set of initial conditions We chose this approach because it is well suited to implementation on digital computers

Numerical radiative transfer in Be disks Deborah Anne Brown,1983

Radiative Transfer in Stellar and Planetary Atmospheres Lucio Crivellari,Sergio Simón-Díaz,María Jesús Arévalo,2020-01-09 Radiative transfer is essential for obtaining information from the spectra of astrophysical objects This volume provides an overview of the physical and mathematical background of radiative transfer and its applications to stellar and planetary atmospheres It covers the phenomenology and physics of early type and late type stars as well as ultra cool dwarf stars and extrasolar planets

Importantly it provides a bridge between classical radiative transfer and stellar atmosphere modelling and novel approaches from both theoretical and computational standpoints With new fields of application and a dramatic improvement in both observational and computational facilities it also discusses the future outlook for the field Chapters are written by eminent researchers from across the astronomical disciplines where radiative transfer is employed Using the most recent observations this is a go to resource for graduate students and researchers in astrophysics

Radiative Transfer in the Atmosphere and Ocean Gary E. Thomas, Knut Stamnes, 2002-01-28 Provides a foundation of the theoretical and practical aspects of radiative transfer for the atmospheric oceanic and environmental sciences

Numerical Regularization for Atmospheric Inverse Problems Adrian Doicu, Thomas Trautmann, Franz Schreier, 2010-07-16 The retrieval problems arising in atmospheric remote sensing belong to the class of the called discrete ill posed problems These problems are unstable under data perturbations and can be solved by numerical regularization methods in which the solution is stabilized by taking additional information into account The goal of this research monograph is to present and analyze numerical algorithms for atmospheric retrieval The book is aimed at physicists and engineers with some background in numerical linear algebra and matrix computations Although there are many practical details in this book for a robust and efficient implementation of all numerical algorithms the reader should consult the literature cited The data model adopted in our analysis is semi stochastic From a practical point of view there are no significant differences between a semi stochastic and a deterministic framework the differences are relevant from a theoretical point of view e.g. in the convergence and convergence rates analysis After an introductory chapter providing the state of the art in passive atmospheric remote sensing Chapter 2 introduces the concept of ill posedness for linear discrete equations To illustrate the difficulties associated with the solution of discrete ill posed problems we consider the temperature retrieval by nadir sounding and analyze the solvability of the discrete equation by using the singular value decomposition of the forward model matrix

Literature 1987, Part 2 U. Esser, H. Hefele, I. Heinrich, W. Hofmann, D. Krahn, V. R. Matas, L. D. Schmadel, G. Zech, 2013-11-11 Astronomy and Astrophysics Abstracts aims to present a comprehensive documentation of the literature concerning all aspects of astronomy astrophysics and their border fields It is devoted to the recording summarizing and indexing of the relevant publications throughout the world Astronomy and Astrophysics Abstracts is prepared by a special department of the Astronomisches Rechen Institut under the auspices of the International Astronomical Union Volume 44 records literature published in 1987 and received before February 15 1988 Some older documents which we received late and which are not surveyed in earlier volumes are included too We acknowledge with thanks contributions of our colleagues all over the world We also express our gratitude to all organizations observatories and publishers which provide us with complimentary copies of their publications Dr Siegfried Böhme retired from his duties as co editor of Astronomy and Astrophysics Abstracts on December 31 1987 Since 1950 he participated in the bibliographic work of the institute He served as a reviewer for the Astronomischer Jahresbericht and

became one of the editors of Astronomy and Astrophysics Abstracts in 1969. After his retirement in 1975 he took care of particularly the Russian literature on a voluntary basis for 12 years. It is a pleasure to thank Siegfried Böhme for his valuable contributions. Starting with Volume 33 all the recording, correction and data processing work was done by means of computers. The recording was done by our technical staff members Ms Helga Ballmann, Ms Christiane Jehn, Ms Monika Kohl, Ms *Fundamentals of Solar Radiation* Lucien Wald, 2021-05-12. The sun radiates a tremendous amount of energy called solar energy or solar radiation which is the main natural source of energy on the Earth by far. Because solar radiation is the almost unique supplier of energy to the Earth it has a primary influence on life and activities on the Earth. The climate is a first example but there are many others such as plant growth or human health or even the design of buildings, the production of energy notably electrical and thermal or even aging materials. This book aims to provide simple answers to anyone who has questions about solar radiation. Its ambition is to help by presenting the fundamental elements of the solar radiation received on the ground. The book includes many examples and numerous illustrations as well as some simple but fairly precise equations to calculate the various elements covered and to reproduce the figures and graphs. The first of the three parts of this book is devoted to the relative geometry between the direction of the sun and an observer on the ground as well as to the solar radiation emitted by the sun and received at the top of the atmosphere. The orbit of the Earth around the sun and the solar declination are described. The concept of time is introduced which is closely linked to the solar cycle and the rotation of the Earth on itself. Equations are given to calculate the solar radiation received on a horizontal or inclined surface located at the top of the atmosphere. The spectral distribution of the extraterrestrial solar radiation is described. The second part of this book addresses how the solar radiation incident at the top of the atmosphere is attenuated and modified in its downward path to the ground. The reflection of the radiation by the ground is presented. The solar radiation received on the ground by a horizontal or inclined collector plane such as a natural slope or a rooftop is discussed as well as its spectral distribution. The variability of the radiation is addressed in relation to the properties of solar radiation estimated from the measurements. The third part deals with direct or indirect measurements of the solar radiation received on the ground over a given integration time: minute, hour, day or month, whether for total radiation or radiation in a spectral range such as ultraviolet (UV) or daylight or photosynthetically active radiation (PAR). It also explains how to check the plausibility of the measurements. *Fundamentals of Solar Radiation* will be a valuable resource to all professionals, engineers, researchers, students and other practitioners that seek an understanding of solar radiation. Computational Heat Transfer Yogesh Jaluria, 2002-11-21. This new edition updated the material by expanding coverage of certain topics, adding new examples and problems, removing outdated material and adding a computer disk which will be included with each book. Professor Jaluria and Torrance have structured a text addressing both finite difference and finite element methods, comparing a number of applicable methods. *The Numerical Solution of Radiative Transfer Problems* David Ellison, 1973. Radiative Transfer in Curved Media K. K. Sen, S. J.

Wilson,1990 Most of the methods described in this book can be used with cosmetic modifications to solve transfer problems of greater complexity All attempts have been made to make the book self contained

Radiation Trapping in Atomic Vapours Andreas F. Molisch,Bernhard P. Oehry,1998 Radiation from spectral lines can be absorbed and re emitted many times in atomic vapours before it reaches the boundaries of the container encasing the vapour This effect is known as radiation trapping It plays an important role practically everywhere where atomic vapours occur e g in spectroscopy in gas lasers in atomic line filters in the determination of atomic lifetimes in measurements of atomic interaction potentials and in electric discharge lamps This book for the first time assembles all the information necessary for a treatment of practical problems emphasizing both physical insights and mathematical methods After an introduction that reviews resonance radiation and collisional processes in atomic vapours physical effects and mathematical methods for various types of problems e g with or without saturation particle diffusion reflecting cell walls etc are explained in detail The last part of the book describes the applications of these methods to a variety of practical problems like cross section measurements or the design of discharge lamps

Thermal Radiation Heat Transfer, Fourth Edition Robert Siegel,2001-12-07 This extensively revised 4th edition provides an up to date comprehensive single source of information on the important subjects in engineering radiative heat transfer It presents the subject in a progressive manner that is excellent for classroom use or self study and also provides an annotated reference to literature and research in the field The foundations and methods for treating radiative heat transfer are developed in detail and the methods are demonstrated and clarified by solving example problems The examples are especially helpful for self study The treatment of spectral band properties of gases has been made current and the methods are described in detail and illustrated with examples The combination of radiation with conduction and or convection has been given more emphasis nad has been merged with results for radiation alone that serve as a limiting case this increases practicality for energy transfer in translucent solids and fluids A comprehensive catalog of configuration factors on the CD that is included with each book provides over 290 factors in algebraic or graphical form Homework problems with answers are given in each chapter and a detailed and carefully worked solution manual is available for instructors

Transactions of the International Astronomical Union Jean-Pierre Swings,2012-12-06 The XIXth General Assembly of the International Astronomical Union was held in New Delhi India from November 19 to 28 1985 It was dedicated to the memory of a former IAU President Professor M K V Bappu who tragically passed away on August 19 1982 On the occasion of the Delhi General Assembly the IAU Minor Planet Center announced that Minor Planet asteroid No 2596 henceforth will carry the name Vainu Sappu The full text of the announcement reads 2596 VAINU BAPPU 1979 KN diameter about 8 kilometers period 5 years 4 months mean distance from the Sun around 450 million kilometers Discovered 1979 May 19 by R M West at the European Southern Observatory Named in memory of Manali f

Light Scattering Reviews 5 Alexander A. Kokhanovsky,2010-08-05 Light scattering by densely packed inhomogeneous media is a particularly ch lencing

optics problem. In most cases only approximate methods are used for the calculations. However, in the case where only a small number of macroscopic scattering particles are in contact clusters or aggregates, it is possible to obtain exact results solving Maxwell's equations. Simulations are possible, however, only for a relatively small number of particles, especially if their sizes are larger than the wavelength of incident light. The first review chapter in Part I of this volume, prepared by Yasuhiko Okada, presents modern numerical techniques used for the simulation of optical characteristics of densely packed groups of spherical particles. In this case, Mie theory cannot provide accurate results because particles are located in the near field of each other and strongly interact. As a matter of fact, Maxwell's equations must be solved not for each particle separately but for the ensemble as a whole in this case. The author describes techniques for the generation of shapes of aggregates. The orientation averaging is performed by a numerical integration with respect to Euler angles. The numerical aspects of various techniques, such as the T-matrix method, discrete dipole approximation, the finite difference time domain method, effective medium theory, and generalized multi-particle Mie solution, are presented. Recent advances in numerical techniques, such as the grouping and adding method and also numerical orientation averaging using a Monte Carlo method, are discussed in great depth.

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