

Selection criterion	Groups				Total	Fraction of stars with $E(B-V) \leq 0.5$
	A	B	C	D		
	Parallax $\text{SNR} \geq 5$		$3 \leq \text{Parallax SNR} < 5$			
	ALLWISE	2MASS	ALLWISE	2MASS		
Cut #0 (initial set)	1484627	1007019	509965	122681	3124292	$\sim 85\%$
Cut #1	3218	891	720	488	5317	$\sim 67\%$
Cut #2	552	162	132	170	1016	$\sim 48\%$
Cut #3	299	78	118	154	649	$\sim 45\%$
Known dLHdC/RCB stars removed	12 <sup>*A</sup>	4 <sup>*B</sup>	1 <sup>*C</sup>	1 <sup>*D</sup>	18	$\sim 55\%$
Total selected	287	74	117	154	631	$\sim 44\%$
Observed	139	43	84	91	357	$\sim 65\%$
New dLHdC/RCB stars found	11	6	9	0	26	$\sim 46\%$

\*A: LHS 137613 LHS 173400 LHS 149930 LHS 175903 V7553 Cyg V533 Cyg ASAS RCB 10

# Selected Exercises In Galactic Astronom

**I. Atanasijevic**



## **Selected Exercises In Galactic Astronom:**

*Selected Exercises in Galactic Astronomy*, 1971      *Selected Exercises in Galactic Astronomy* I. Atanasijevic, 2012-12-06  
Galactic Astronomy is and for some time to come will certainly remain one of the more important items in the astronomy curricula. There are at present a number of excellent text books and monographs covering the whole subject or certain of its aspects. It seems however that there is not a single book where at least some of the more important problems dealt with in text books and university courses would be presented in the form of laboratory exercises. This short series of exercises represents an attempt to fill this gap. What is in general the aim of such exercises? As to this point the author fully agrees with Prof. Minnaert's opinion that no natural science should ever be taught without practical work. Students in astronomy should have regular exercises not so much to teach them observing skill but mainly to bring before their eyes the reality of the concepts introduced during the lectures. This applies to Galactic Astronomy as well as to any other branch of our science.

*Selected exercises in galactic astronomy* Ivan M. Atanasijević, 1973      ***Exercises in Astronomy*** J. Kleczek, 2012-12-06  
*Revised and Extended Edition of Practical Work in Elementary Astronomy* by M. G. J. Minnaert      ***Stellar Paths*** P. Kamp, 2012-12-06  
This is the latest effort in a sequence of presentations begun in 1949 with a series of lectures on long focus photographic astrometry given by the author as Fulbright professor in Paris at the invitation by the late H. Mineur, at that time Director of the Institut d'Astrophysique. These earlier lectures were published as a series of review articles in *Popular Astronomy* 1951 and appeared both as *Contributions de l'Institut d'Astrophysique Serie A No 81* and as reprint No 75 of *Sproul Observatory*. A more elaborate presentation was given in 1963 in *Stars and Stellar Systems* which was followed by *Principles of Astrometry* 1967 W. H. Freeman Co. During the second half of 1974 again under Fulbright auspices at the invitation of P. K. Sin, I lectured at the Astronomical Institute in Amsterdam followed by a short course in May/June 1978 at the invitation of E. P. J. van den Heuvel. I gave a more extensive course at the Institut d'Astrophysique at the invitation of J. C. Pecker of the College de France and of J. Audouze, Director of the LA P. Both in Amsterdam and in Paris I had presented occasional astrometric topics at various times. The opportunity to lecture in France and in Holland has facilitated, influenced and improved the organization and contents of the presentations on the subject of long focus photographic astrometry.

***Cosmic Plasma*** H. Alfvén, 2012-12-06  
The general background of this monograph and the aim of it is described in detail in Chapter I. As stated in 1.7 it is written according to the principle that when rigour appears to conflict with simplicity, simplicity is given preference, which means that it is intended for a rather broad public. Not only graduate students but also advanced undergraduates should be able to understand at least most of it. This monograph is the result of many years of inspiring discussions with a number of colleagues for which I want to thank them very much. Especially I should mention the groups in Stockholm and La Jolla in Stockholm. Dr. Carl Gunne, Flilthammar and many of his collaborators including Drs. Lars Block, Per Carlqvist, Lennart Lindberg, Michael Raadu, Staffan Torven, Miroslav Babic and Itlgvar Axniis and further Drs. Bo

Lehnert and Bjorn Bonnevier all at the Royal Institute of Technology Of other colleagues in Sweden I should mention Dr Bertel Laurent Stockholm University Dr Aina Elvius The Stockholm Observatory and Dr Bengt Hultqvist Kiruna In La Jolla my thanks go first of all to Dr Gustaf Arrhenius who once invited me to La Jolla which was the start of a most interesting collaboration further to Dr W B

**Multivariate Data Analysis** Fionn Murtagh, Andre Heck, 2012-12-06 Interest in statistical methodology is increasing so rapidly in the astronomical community that accessible introductory material in this area is long overdue This book fills the gap by providing a presentation of the most useful techniques in multivariate statistics A wide ranging annotated set of general and astronomical bibliographic references follows each chapter providing valuable entry points for research workers in all astronomical sub disciplines Although the applications considered focus on astronomy the algorithms used can be applied to similar problems in other branches of science Fortran programs are provided for many of the methods described

**Physical Processes in Red Giants** I. Iben, Alvio Renzini, 2012-12-06 In recent years it has become clear that the red giant phase is one of the most dramatic periods in a star's life when all of its parts become involved in ways that have both direct and indirect observational consequences This is most particularly true of low and intermediate mass stars during the second ascent of the giant branch Such stars bring to their surfaces products of nucleosynthesis currently taking place in their deep interiors they pulsate as Mira variables develop extended outward flowing atmospheres that may exhibit maser properties and shed great quantities of matter sometimes highly processed into the inter stellar medium The manner in which processed matter is brought to the surface is far from being completely explained and the precise mechanism or mechanisms whereby matter is ejected from the stellar surface whether by deposition of Alfvén waves radiation pressure on grains or as a consequence of some large scale envelope instability has yet to be elucidated to everyone's satisfaction The purpose of the second workshop in Astrophysics organized by the Advanced School of Astronomy was to bring together experts on all the physical processes occurring in red giants in an effort to emphasize the interrelatedness of these individual processes and to encourage a dialogue among experts that might serve to initiate a synthesis or at least sharpen our understanding of the most important problems to address in the future

**Astrophysics from Spacelab** P.L. Bernacca, Remo Ruffini, 2012-12-06 A meeting on Astrophysics from Spacelab was held at the International Centre for Theoretical Physics Trieste in the Autumn of 1976 Scope of the meeting was to bring to the attention of an increasing number of physicists and astrophysicists including scientists from developing countries the new facilities made available by the combination of the Shuttle and the Spacelab programmes This book starts from that meeting and includes together with reports presented in Trieste duly updated a few additional reviews on selected topics In the first part D J Shapland and G Giampalmo The Shuttle and the Spacelab present the design and the programmatic data of these advanced transportation systems and orbital laboratories Vittorio Manno introduces the scientific programmes coordinated and led to execution by the European Space Agency ESA Programmes in Astronomy and Astrophysics J D

Rosendha 1 liThe NASA Programmes in Astronomy and Astrophysics summarizes the activities in solar physics high energy astrophysics and astronomy planned in the United States of America by the National Aeronautics and Space Administration as well as the expected use of the space shuttle and spacelab in their first year of operation The Solar Chromosphere and Corona: Quiet Sun R.G. Athay,2012-12-06 The widespread tendency in solar physics to divide the solar atmosphere into separate layers and to distinguish phenomena of solar activity from phenomena of the quiet Sun emphasizes the wide ranging diversity of physical conditions and events occurring in the solar atmosphere This diversity spans the range from a neutral essentially quiescent atmosphere to a highly ionized violently convective atmosphere from a domain in which magnetic field effects are unimportant to a domain in which the magnetic pressure exceeds the gas pressure and from a domain in which the particle motions are Maxwellian to a domain in which an appreciable fraction of the particles is accelerated to relativistic energies It is now widely recognized that the chromosphere and corona have a common origin in the mechanical energy flux generated in the hydrogen convection zone lying beneath the photosphere Furthermore magnetic field phenomena appear to be as vital to the structure of the quiet Sun as to the active Sun For these reasons it appears desirable to present a unified treatment of the entire solar atmosphere both active and quiet in a single volume On the other hand such a treatise must be very long if it is to avoid being superficial and it is very difficult for a single author to write authoritatively on such a wide range of topics **Jacobi Dynamics** V.I. Ferronsky,S.A. Denisik,S.V. Ferronsky,2012-12-06 This book sets forth and builds upon the fundamentals of the dynamics of natural systems in formulating the problem presented by Jacobi in his famous lecture series Vorlesungen tiber Dynamik Jacobi 1884 In the dynamics of systems described by models of discrete and continuous media the many body problem is usually solved in some approximation or the behaviour of the medium is studied at each point of the space it occupies Such an approach requires the system of equations of motion to be written in terms of space co ordinates and velocities in which case the requirements of an internal observer for a detailed description of the processes are satisfied In the dynamics discussed here we study the time behaviour of the fundamental integral characteristics of the physical system i e the Jacobi function moment of inertia and energy potential kinetic and total which are functions of mass density distribution and the structure of a system This approach satisfies the requirements of an external observer It is designed to solve the problem of global dynamics and the evolution of natural systems in which the motion of the system s individual elements written in space co ordinates and velocities is of no interest It is important to note that an integral approach is made to internal and external interactions of a system which results in radiation and absorption of energy This effect constitutes the basic physical content of global dynamics and the evolution of natural systems **Radio Recombination Lines** P.A. Shaver,2012-12-06 Since their first detection 15 years ago radio recombination lines from several elements have been observed in a wide variety of objects including HII regions planetary nebulae molecular clouds the diffuse interstellar medium and recently other galaxies The observations span almost the entire

range from 0.1 to 100 GHz and employ both single dish and aperture synthesis techniques. The theory of radio recombination lines has also advanced strongly to the point where it is perhaps one of the best understood in astrophysics. In a parallel development it has become possible over the last decade to study these same highly excited atoms in the laboratory; this work provides further confirmation of the theoretical framework. However, there has been continuing controversy over the astrophysical interpretation of radio recombination line observations, especially regarding the role of stimulated emission. A workshop was held in Ottawa on 24-25 August 1979, bringing together many of the active scientists to review the field and discuss these questions of interpretation. A broad consensus has emerged: the subtleties of the line formation process are understood and the conditions under which reliable information can easily be extracted from the line measurements are known. It thus appears likely that the emphasis will shift increasingly from the study of the line phenomenon itself to further application in other areas of astrophysics, ranging from physical processes in plasmas (temperatures, densities, ionization structure) to the large scale properties of our galaxy (abundances, kinematics, structure) and studies of extragalactic systems.

**Planetary Nebulae** Stuart R. Pottasch, 2012-12-06. The purpose of this book is to give a detailed description of the planetary nebulae, including the relevant astronomical observations and their interpretation. Considerable attention is given to the evolution of these objects as well as to their physical characteristics. It is hoped that the book be useful to both advanced research workers and to students with some background in astronomy. In this regard, the book should serve as a text as well as a reference work. The many tables included are expected to be useful for both purposes. The references are generally not included in the text except for historical purposes in an effort to improve readability. References are given at the end of each chapter together with sufficient text to describe their content. No attempt has been made to make the list of references complete; on the contrary, it has generally been limited to the most recent literature on the subject, which in turn refers to earlier research. Again, exceptions have been made for references of historical interest. *Language of the Stars*

Zdenek Kopal, 2012-12-06. *Eclipsing Variables: What They can Tell Us and What We can do with Them*. The aim of the present book will be to provide an introduction to the interpretation of the observed light changes of eclipsing binary stars and their analysis for the elements of the respective systems. Whenever we study the properties of any celestial body, be it a planet or a star, all information we wish to gain can reach us through two different channels: their gravitational attraction and their light. Gravitational interaction between our Earth and its celestial neighbours is, however, measurable only at distances of the order of the dimensions of our solar system, and the only means of communication with the realm of the stars are their nimble footed photons, reaching us with appropriate time lag across the intervening gaps of space. As long as a star is single and emits constant light, it does not constitute a very revealing source of information. A spectrometry of its light can disclose to be sure the temperature, colour, or ionization of the star's semi-transparent outer layers, their chemical composition, and prevalent pressure. Through Stark effect or magnetic field Zeeman effect, it can disclose even some information about its

absolute luminosity or rate of spin It cannot however tell us anything about what we should like to know most namely the mass or size i e density of the respective configuration its absolute dimensions or its internal structure     The Magnetospheres of the Earth and Jupiter V. Formisano,2012-12-06 The Symposium The Magnetospheres of the Earth and Jupiter May 28th June 1st 1974 in Frascati was organized by the Laboratorio Plasma Spazio of the CNR on the occasion of its moving to the Frascati area The main theoretical topic was to be covered by N Brice but he died on January 31st in a plane crash at Pago Pago Samoa It seemed appropriate to all of us to honor Neil Brice by renaming the meeting Neil Brice Memorial Symposium The Symposium covered problems of magnetospheric dynamics of both the Earth and Jupiter with emphasis on the comparison between the two objects The collaboration of American scientists participating in the recent NASA planetary missions has made it possible to have new important scientific results presented to and discussed by the scientific community Of the many people who helped to make the meeting such a success very special thanks goes to Prof C Kennel of U C L A whose contribution to the organization of this symposium has been extremely important     *Correlated Interplanetary and Magnetospheric Observations* D.E. Page,2012-12-06 The 1969 ESLAB symposium Intercorrelated Satellite Observations Related to Solar Events was held at a time when the importance of bringing together measurements made simultaneously in different regions of space was beginning to be appreciated To day it is universally accepted that the major experimental steps forward in understanding the physics of the Sun Earth relationships are likely to be made through pre planned correlated satellite studies Such considerations have led to the organisation of the International Magnetospheric Study and the joint ESRO NASA International Magnetospheric Explorer Mother Daughter Heliocentric mission The seventh ESLAB symposium was planned as a follow up to that of 1969 with the aim of deriving maximum benefit from those spacecraft which through good fortune found themselves simultaneously operating in different regions of the magnetosphere and interplanetary space ESRO had launched in early 1972 its HEOS 2 satellite to investigate fields and particles in the unexplored region far above the North pole of the Earth and it became clear that the interesting new results arriving from that mission could profitably be linked with those from various American and the U S S R PROGNOZ satellites The book follows the order of the symposium unfortunately the PROGNOZ contribution did not materialise concentrating through both review lectures and of new experimental results on the nature of the boundaries between the interplanetary medium and the magnetosphere and the interaction of each region on the other     Ionospheric Effects of Solar Flares Hermine Vloemans,2012-12-06 Sudden Ionospheric Disturbances resulting from an interaction of the Solar Flare radiation with the constituents of the upper atmosphere constitute one of the three major aspects of ground level monitoring of solar flares the other two being optical observations of flares and the observations of solar bursts in radio wavelengths SIDs therefore form a major part of flare monitoring programme in many observatories Unlike the other two however the ionospheric effects of flares provide one major additional source of interest the reaction of the ionospheric plasma to an

impulsive ionization The high atmosphere provides a low pressure laboratory without walls in which a host of reactions occur between electrons ions and neutral particles The resulting products and their distributions may bear no resemblance to those of the primary neutral constituents or their direct ionization products The variations with the time of the day with season and with solar activity that form the bulk of the ionospheric measurements are too slow to allow any insight into the nature of these ionospheric reactions whose lifetimes are often very short The relaxation time of the ionospheric ionization is only a few minutes or fraction of a minute in the lower ionosphere and in the E region and is about 30 min to an hour at 300 km The flares provide a sudden short impulse comparable to these time scales

*Atmospheres of Earth and the Planets* Billy McCormac, 2012-12-06 This book contains the lectures presented at the Summer Advanced Study Institute Physics and Chemistry of Atmospheres which was held at the University of Liege Belgium during the period July 29 August 9 1974 One hundred nineteen persons from eleven different countries attended the Institute The authors and publisher have made a special effort for rapid publication of an up to date status of the physics and chemistry of the atmospheres of Earth and the planets which is an ever changing area Special thanks are due to the lecturers for their diligent preparation and excellent presentations The individual lectures and the published papers were deliberately limited the authors cooperation in conforming to these specifications is greatly appreciated The contents of the book are organized by subject area rather than in the order in which papers were presented during the Institute Many thanks are due to Drs Alv Egeland Donald M Hunten Gunther Lange Hesse Marcel Nicolet Harold I Schiff Lance Thomas Alister Vallance Jones Richard Wayne and Gilbert Weill who served as session chairmen during the Institute and contributed greatly to its success by skillfully directing the discussion period in a stimulating manner after each lecture Many persons contributed to the success of the Institute Drs Alv Egeland Donald M Hunten Gunther Lange Hesse Marcel Nicolet Harold I Schiff Erwin R Schmerling Lance Thomas Alister Vallance Jones Richard Wayne and Gilbert Weill were especially helpful in preparing the technical program

**Neutron Stars, Black Holes and Binary X-Ray Sources** H. Gursky, Remo Ruffini, 2013-04-17 This book contains a set of articles based on a session of the annual meeting of the American Association for the Advancement of Science held in San Francisco in February 1974 The reason for the meeting arose from the need to communicate to the largest possible scientific community the dramatic advances which have been made in recent years in the understanding of collapsed objects neutron stars and black holes Thanks to an unprecedented resonance between X ray y ray radio and optical astronomy and important new theoretical developments in relativistic astrophysics a new deep understanding has been acquired of the physical processes occurring in the late stages of evolution of stars This knowledge may be one of the greatest conquests of man's understanding of nature in this century This book aims to give an essential and up to date view in this field The analysis of the physics and astrophysics of neutron stars and black holes is here attacked from both theoretical and experimental points of view In the experimental field we range from the reviews and catalogues of galactic X ray sources R Gursky and E Schreier



and pulsars E Groth to the observations of the optical counter part of X ray sources P Boynton to finally the recently discovered gamma ray bursts I Strong and pulse astronomy R B Partridge      Mapping of the Moon Zdenek Kopal, R.W. Carder, 2013-04-17 The aim of the present book has been to provide an outline the first of its kind of the history of the human efforts to map the topography of the surface of our satellite from the days of pre telescopic astronomy up to the present These efforts commenced modestly at the time when the unaided eye was still the only tool at the disposal of men interested in the face of our satellite and were continued since for more than three centuries by a small band of devoted friends of the Moon in several countries Many of these were amateur astronomers and almost all were amateur cartographers though some highly skilled in their art The reader interested in the history of lunar mapping between 1600 and 1960 will find its outline in the first chapter of this book and can follow the way in which the leadership in the mapping of the Moon the cradle of which stood in Italy passed successively to France Germany and eventually to the United States All efforts described in this chapter were wholly superseded by subsequent developments since 1960 largely motivated by logistic needs of a grand effort which culminated with repeated manned landings on the Moon between 1969 1972 a feat which will remain for ever one of the glories of our century

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