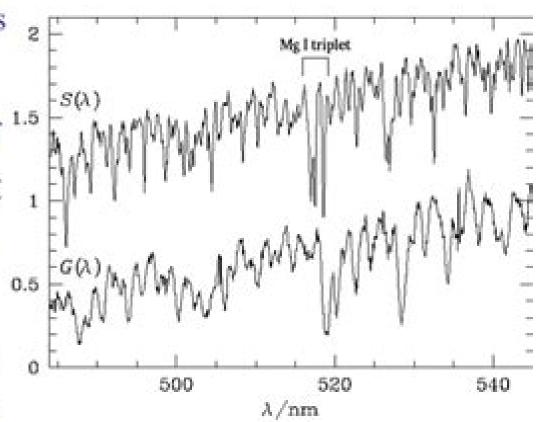
The kinematics of stars in E galaxies

Stars in E galaxies do not follow ordered motions, but most of their kinetic energy is in the form of random motions. Moreover, the more luminous an E, the larger the velocity dispersion (which, as we have seen can be used to derive distances to E).

To measure the orbital velocities of stars within galaxies is a difficult task. Absorption lines are used, which are the result of the light of all stars. Each star emits a spectrum which is Doppler shifted in wavelength according to its motion. This orbital motion makes the resulting spectral lines be wider than those of an individual star.



Photometry Kinematics And Dynamics Of Galaxies

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for which identification atlases are already in print and shows the most important objects including many thousands of individual stars and interstellar objects It is unique in its coverage and format and provides a source of these fundamental data that will be used for many years Researchers students and even amateur astronomers will be able to use the Atlas to identify and study the various components of the nearly 30 important galaxies covered by the Atlas The objects identified on the more than 200 charts include variable stars globular star clusters open star clusters stellar associations emission regions supernova remnants planetary nebulae and dust clouds Each galaxy is accompanied by an extensive bibliography

Morphological and Physical Classification of Galaxies G. Longo, Massimo Capaccioli, G. Busarello, 1992-05-31 The morphological scheme devised by Hubble and followers to classify galaxies has proven over many decades to be guite effective in directing our quest for the fundamental pa rameters describing the extragalactic manifold This statement is however far more true for spirals than for ellipticals Echoing the concluding remarks in Scott Tremaine s sum mary talk at the Princeton meeting on Structure and Dynamics of Elliptical Galaxies the Hubble classification of spirals is useful because many properties of spirals gas con tent spiral arm morphology bulge prominence etc all correlate with Hubble time By contrast almost nothing correlates with the elliptical Hubble sequence El to E7 During the last few years much effort has been put into the search for a more meaningful classification of ellipticals than Hubble's Concomitantly forwarded by some provocative conjectures by R Michard the classical question of whether E galaxies form a physically homogeneous family has been brushed up once more Results of these and other parallel studies look rather promising and point to suture part of the dichotomy between ellipticals and disk galaxies which had become popular in the early eighties owing to dynamical arguments At the same time it appears more and more clear that besides the usual genetic varieties of galaxies products of environmental evolution must also be contemplated in building our modern picture of the reign of galaxies The above The World of Galaxies Harold G. Corwin, Lucette Bottinelli, 2012-12-06 From 12 considerations prompted us to solicit Prof April to 14 April 1988 120 of Gerard and Antoinette de Vaucouleurs s friends and colleagues gathered at the Institut d Astrophysique in Paris to cel ebrate Gerard s 70th birthday and his remarkable career in Astronomy The gathering also honored the memory of Antoinette who died 29 August 1987 after a long illness and her own no less remarkable career This volume collects the 24 invited review papers and the 60 contributed poster papers presented at the meeting Gerard de Vaucouleurs Gerard de Vaucouleurs was born on 25 April 1918 in Paris where he spent his boyhood He became an active amateur astronomer in the early 1930 s making extensive observations of Mars Jupiter and variable stars including the bright supernova of 1937 in IC 4182 He also began life long interests in astronomical photography and galaxy cataloguing during this period In 1939 he met the director of the Paris transport system and an equally avid amateur astronomer Julien Peridier De Vaucouleurs worked at Peridier's private observatory at Le Houga in southwestern France on and off throughout the next decade His undergraduate work was in mathematics astronomy and experimental physics this combined with his interest in

observational astronomy formed his life long empirical approach to science After spending 18 months in the French army early in 1939 41 Gerard returned to the Peridier observatory and then in 1943 to his studies at the Sorbonne where he met Antoinette Active Galactic Nuclei John Edward Dyson, 1985 Formation and Evolution of Galaxy Bulges (IAU S245) International Astronomical Union. Symposium, 2008-09-04 Bulges lie at the heart of most galaxies the building blocks of our universe With a selection of reviews and topical presentations IAU Symposium 245 provides an up to date overview of our knowledge on galaxy bulges and a concise introduction to all current research on the subject The structure dynamics and stellar populations of galaxy bulges both near and far are analysed through state of the art observations. The leading models for the formation and evolution of galaxy bulges are described in detail and the constraints observations put on these are dissected Particular emphasis is placed on exploring evidence for both hierarchical merging and secular processes Special attention is also devoted to disentangling the complex web relating galaxy bulges and central supermassive black holes and on the lessons learned from our exquisite knowledge of the bulge of our own Milky Way This volume is the best one stop reference on galaxy bulges currently available The de Vaucouleurs Atlas of Galaxies Ronald J. Buta, Harold G. Corwin, Stephen C. Odewahn, 2007-03-08 New illustrated atlas on modern galaxy classification for astronomy researchers students and amateurs Third Asian-Pacific Regional Meeting of the International Astronomical Union M. Kitamura, E. Budding, 2012-12-06 M KITAMURA Tokyo Astronomical Observatory Japan and E BUDDING Carter Observatory Wellington New Zealand The Third Asian Pacific Regional Meeting of the International Astronomical Union was held from 30 September to 5 October 1984 at the Kyoto International Conference Hall Kyoto Japan under the auspices of the Union and the Astronomical Society of Japan with Kyoto University as host Three hundred and twenty seven astronomers from twenty two countries participated at the meeting and more than two hundred papers were presented The aim of the meeting was not only to promote scientific developments and cooperation but also to offer a chance for all participants to become acquainted with major astronomical projects of the Asian Pacific Region Therefore two new sessions of A View of Asian Pacific Astronomy and Astronomical Education in the Asian Pacific Region which had not been undertaken in the previous two Regional Meetings were arranged as a first trial besides the other ordinary scientific sessions The Scientific Organizing Committee consisted of D C Morton chairman R N Manchester S M Gong K J Feng C S Shen J C Bhattacharyya G Swa B Hidayat H M K AI Nairniy H S Yun J B Hearnshaw S C Wolff I Ka rup waguchi M Kitamura M Morimoto M Oda and J P Swings IAU ex officio and the Local Organizing Committee of T Kogure chairman T Ishizawa M Saite R Hirata S Inagaki E Hiei M Kitamura B Takase N Kaifu H Maehara Y Osaki and A Yamasaki Star Formation in Galaxies ,1987

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