# Mass-Transfer Induced Activity in Galaxies

# **Mass Transfer Induced Activity In Galaxies**

Lodewijk Woltjer, G. Setti, J.-P. Swings

# **Mass Transfer Induced Activity In Galaxies:**

Mass-transfer Induced Activity in Galaxies Isaac Shlosman,1994 Mass-Transfer Induced Activity in Galaxies Isaac Shlosman,2003-10-30 How gas flows and starbursts light up active galaxies Secular Evolution of Galaxies Jesús Falcón-Barroso, Johan H. Knapen,2012 The formation and evolution of galaxies is one of the most important topics in modern astrophysics Secular evolution refers to the relatively slow dynamical evolution due to internal processes induced by a galaxy s spiral arms bars galactic winds black holes and dark matter haloes It plays an important role in the evolution of spiral galaxies with major consequences for galactic bulges the transfer of angular momentum and the distribution of a galaxy s constituent stars gas and dust This internal evolution is in turn the key to understanding and testing cosmological models of galaxy formation and evolution Based on the twenty third Winter School of the Canary Islands Institute of Astrophysics this volume presents reviews from nine world renowned experts on the observational and theoretical research into secular processes and what these processes can tell us about the structure and formation of galaxies The volume provides a firm grounding for graduate students and early career researchers working on galactic dynamics and galaxy evolution

Galaxies and their Masks David L. Block, Kenneth C. Freeman, Ivânio Puerari, 2010-11-12 Freeman Fellow of the Royal The Central Regions of the Galaxy and Galaxies International Astronomical Union. Symposium, 1998-08-31 Society Proceedings of the August 1997 symposium One hundred and ninety four contributions present comparative studies on the data of the Milky Way and central regions of nearby galaxies Included is information on galactic bulges galactic center star clusters star formation starbursts neutral ISM in the galactic center molecular gas in the nuclei of galaxies gas dynamics in the galactic center the central parsecs of the milky way magnetic and high energy phenomena black holes in galaxies black hole in the galactic center and black hole powering of AGN and jets A sampling of topics diffraction limited IR speckle masking observations of the central regions of Seyfert galaxies the stellar content of the Quintuplet cluster and the structural characteristics of spiral bulges Annotation copyrighted by Book News Inc Portland OR **Reports on Astronomy** Immo Appenzeller, 2012-12-06 IAU Transactions are published as a volume corresponding to each General Assembly Volume A is produced prior to the Assembly and contains Reports on Astronomy prepared by each Commission President The intention is to summarize the astronomical results that have affected the work of the Commission since the production of the previous Reports up to a time which is about one year prior to the General Assembly Volume B is produced after the Assembly and contains accounts of Commission Meetings which were held together with other material The reports included in the present volume range from outline summaries to lengthy compilations and references Most reports are in English The Interplay Between Massive Star Formation, the ISM and Galaxy Evolution D. Kunth, 1996 The Shapes of Galaxies and **Their Dark Halos** Priyamvada Natarajan, 2002 This book constitutes the proceedings of a very topical workshop aimed at understanding the shapes of the baryonic and dark matter components of galaxies Several groups presented their recent

results from observations and numerical N body simulations 
The Interstellar Medium in Galaxies J.M. van der Hulst, 2013-12-01 It has been more than five decades ago that Henk van de Hulst predicted the observability of the 21 cm line of neutral hydrogen HI Since then use of the 21 cm line has greatly improved our knowledge in many fields and has been used for galactic structure studies studies of the interstellar medium ISM in the Milky Way and other galaxies studies of the mass distribution of the Milky Way and other galaxies studies of spiral struc ture studies of high velocity gas in the Milky Way and other galaxies for measuring distances using the Tully Fisher relation etc Regarding studies of the ISM there have been a number of instrumental developments over the past decade large CCD s became available on optical telescopes radio synthesis offered sensitive imaging capabilities not only in the classical 21 cm HI line but also in the mm transitions of CO and other molecules and X ray imaging capabilities became available to measure the hot component of the ISM These developments meant that Milky Way was no longer the exclusive playground for ISM studies and that by reaching out to other galaxies astronomy had gained the advantage of having an outside view though at the expense of giving up some linear resolution Studies of the ISM in other galaxies are intimately connected to studies of the ISM in the Milky Way

Morphology, Kinematics and Star Formation Across the Hubble Sequence of Galaxies Santiago Erroz-Ferrer, 2015-11-21 This thesis discusses the evolution of galaxies through the study of the morphology kinematics and star formation properties of a sample of nearby galaxies The main body of the thesis describes the kinematic observations with the GHaFAS Fabry Perot instrument on the William Herschel Telescope of a sample of 29 spiral galaxies. The work is closely related to the Spitzer Survey of Stellar Structure in Galaxies and uses the mid infrared data of that survey to determine key parameters of the galaxies studied From these data important results are obtained on streaming and other non circular motions in galaxies on the distribution and rates of star formation and on how correlations of these parameters and of the rotation curve shape with basic galaxy parameters yield clues on the evolutionary processes taking place in disk galaxies Laboratory **Directed Research and Development** Lawrence Livermore National Laboratory, 1992 Science with Large Millimetre <u>Arrays</u> Peter Shaver, 2013-06-29 The next major step in millimetre astronomy and one of the highest priority items in radio astronomy today is a large millimetre array with a collecting area 2 of up to 10 000 m A project of this scale will almost certainly require international collaboration at least within Europe and possibly with other major partners elsewhere In order to establish a focal point for this project within Europe a study has been undertaken by the Institut de Radio Astronomie Mil Ii met rique IRAM the European Southern Observatory ESO The Onsala Space Observatory OSO and The Netherlands Foundation for Research in Astronomy NFRA In the context of this project a workshop attended by some 100 participants was held at ESO Garching on December 11 13 1995 to discuss the scientific advances such an array will make possible Throughout the three days of the workshop the strong enthusiasm for the concept of a large millimetre array in the southern hemisphere the Large South ern Array or LSA was obvious and it became clear that such a facility would have a

profound impact on almost all areas of observational astrophysics It was particularly clear that since their main science drivers cosmology and the origins of galaxies stars and planets are the same and their angular resolutions and sensitivities similar the LSA and the VLT would strongly complement each other Starbursts Richard de Grijs, Rosa M. González Delgado, 2006-05-11 Starbursts are important features of early galaxy evolution Many of the distant high redshift galaxies we are able to detect are in a starbursting phase often apparently provoked by a violent gravitational interaction with another galaxy In fact if we did not know that major starbursts existed these conference proceedings testify that we would indeed have difficulties explaining the key properties of the Universe These conference proceedings cover starbursts from the small scale star forming regions in nearby galaxies to galaxy wide events at high redshifts one of the major themes of the conference proved to be scalability i e can we scale up the small scale events to describe the physics on larger scales The key outcome of this meeting and these proceedings is a resounding yes to this fundamental yet profound question The enhanced synergy facilitated by the collaboration among observers using cutting edge ground and space based facilities theorists and modellers has made these proceedings a true reflection of the state of the art in this very rapidly evolving field Ouasars, Agns, and Related Research Across 2000 Lodewijk Woltjer, G. Setti, J.-P. Swings, 2001-07-17 The proceedings of the international conference held in honor or Lodewijk Woltjer focus on the physics origin and space distribution of AGNs and quasars and their relationship to the environment a very wide subject which has attracted much of Lodewijk Wolter's research interest A number of lectures were also dedicated to reviewing the recent observational advances and those that may be attained by the introduction of new and powerful astronomical instrumentation both from the ground and from space in recognition of the central role played by L Woltjer in the promotion of the ESO VLT and of his involvement in shaping ESA s space programme Horizon 2000 <u>Violent Star Formation</u> Guillermo Tenorio-Tagle, Instituto de Astrofísica de Canarias, Royal Greenwich Observatory, 1994-09-22 A valuable overview and a timely update on all aspects of violent star formation in a host of objects for graduate students and researchers across a broad range of research interests

Astronomie Extragalactique Dans L'infrarouge Gary Allan Mamon, Xuan Thuan Trinh, J. Thanh Van Tran, 1997 The Interplay Among Black Holes, Stars and ISM in Galactic Nuclei (IAU S222) International Astronomical Union.

Symposium, 2004 How massive are the largest and smallest nuclear black holes in galaxies Why are the masses of nuclear black holes proportional to those of their host galaxy bulges How is nuclear activity triggered What are the observational signatures of such processes What are the connections between the active nucleus stars and interstellar medium in galaxies Answers to these questions are addressed in this book which presents a compilation of 191 works covering recent observations from X rays to radio wavelengths as well as theoretical modeling of accretion disks stellar populations and galaxy and black hole evolution This volume presents the nuclear activity as a phase in the life of a galaxy which is intimately connected to the evolution of its stars and interstellar medium It brings together recent developments in topics covering

most aspects of galaxy evolution and is a valuable resource for astronomers and graduate students working in extragalactic Penetrating Bars through Masks of Cosmic Dust David L. Block, Ivânio Puerari, K. C. Freeman, R. Groess, Elizabeth K. Block, 2013-04-17 THE EDITORS DAVID L BLOCK AND KENNETH C FREEMAN SOC CO CHAIRS IVANIO PUERARI ROBERT GROESS AND LIZ K BLOCK 1 Harvard College Observatory 1958 The past century has truly brought about an explosive period of growth and discovery for the physical sciences as a whole and for astronomy in particular Galaxy morphology has reached a renaissance The year 1958 The date October 1 The venue Harvard College Observatory The lecturer Walter Baade With amazing foresight Baade penned these words Young stars supergiants and so on make a terrific splash lots of light The total mass of these can be very small compared to the total mass of the system Dr Layzer then asked the key question the discussion raises the point of what this classification would look like if you were to ignore completely all the Population I and just focus attention on the Population II We stand on the shoulders of giants The great observer E E Barnard in his pioneering efforts to photograph the Milky Way devoted the major part of his life to identifying and numbering dusty holes and dust lanes in our Milky Way No one could have dreamt that the pervasiveness of these cosmic dust masks not only in our Galaxy but also in galaxies at high redshift is so great that their penetration is truly one of the pioneering challenges from both space borne telescopes and from the ground Stellar Astrophysics for the Local Group A. Aparicio, A. Herrero, F. Sánchez, 1998-06-13 A review of the new subject of extragalactic stellar astrophysics for both graduate students and researchers working in astrophysics Starbursts Triggers, Nature, and Evolution Bruno Guiderdoni, 2013-11-11 Starbursts are regions of unusually rapid star formation often located in the central parts of galaxies They differ from more normal regions of star formation in terms of the throughput of mass and the rapidity with which the gas is consumed In the last twenty years extensive observational data at most wavelengths have become available on starbursts but many important issues remain to be addressed observationally as well as theoretically How are strong episodes of star formation triggered What is the quantity of gas converted into stars during bursts What is the initial mass function of stars in these events How does the feedback from stars influence the interstellar medium and self regulate star formation What is the subsequent chemical and photometric evolution How do starbursts rule the formation and evolution of galaxies In recent years many observational data at different wavelengths optical radio infrared X ray have become available However these observations are still fragmentary in the sense that different classes of objects have been observed in different ways and the coverage is not consistently deep or complete As a consequence an overall observational picture of starburst galaxies is missing and theoretical understanding and modelling have remained highly tentative The purpose of the school Starbursts Triggers Nature and Evolution was to gather theorists and observers with complementary approaches to the starburst phenomenon in order to summarize the state of the art of the observations and models emphasizing the consistency of the various viewpoints

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